

# A Compact Integrated Raman Spectrometer (CIRS)

**Capability**



**Deployment**

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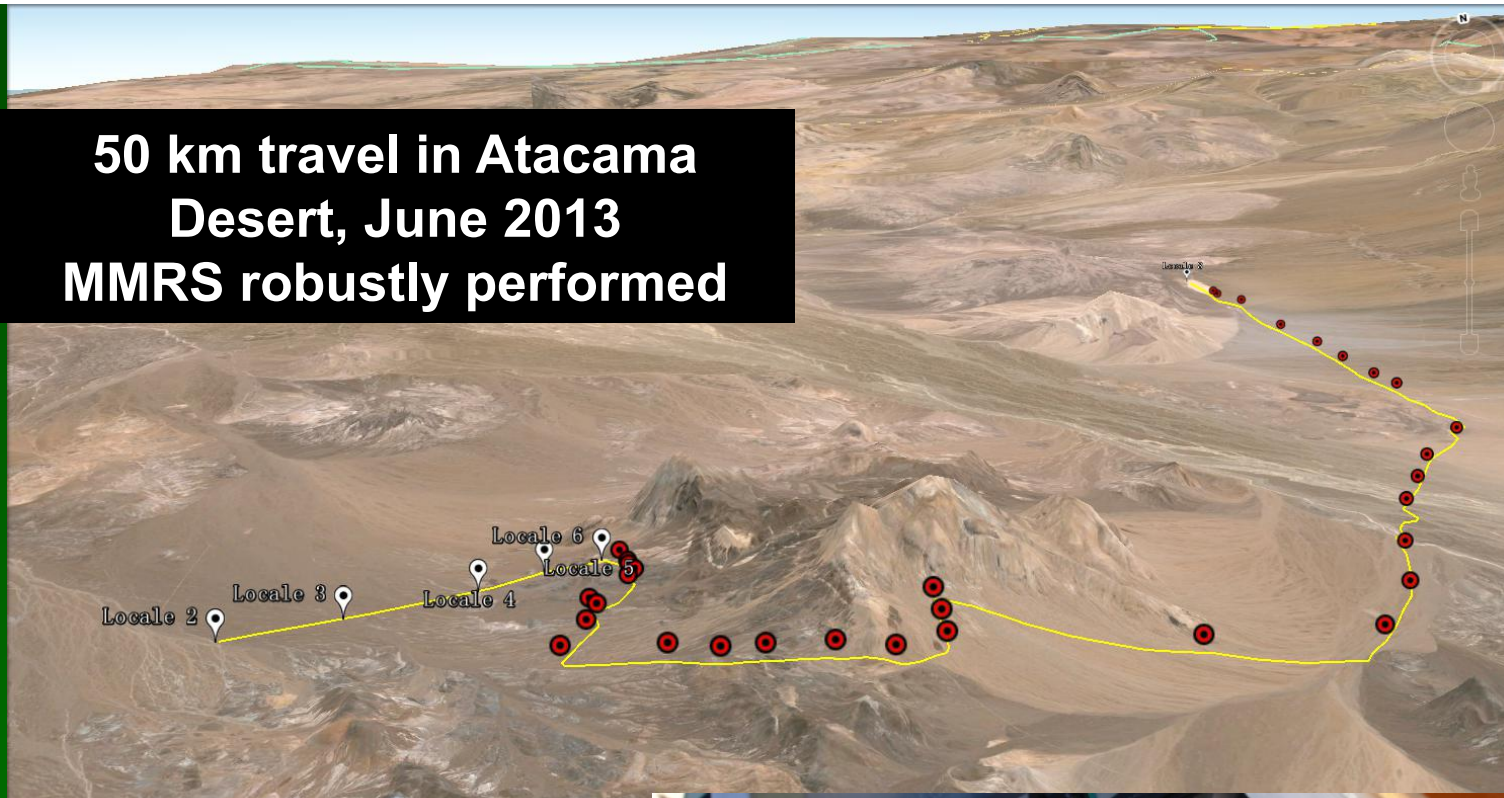
*Technology Workshop for Discovery*  
*Washington DC, April 9, 2014*

# Maturity in Science & Engineering

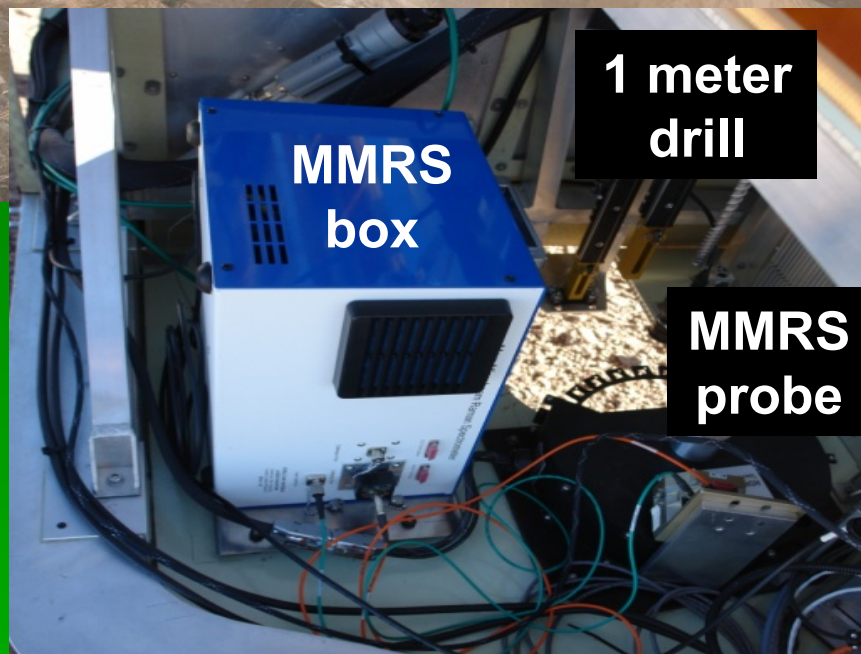
- Mars Microbeam Raman Spectrometer (MMRS) started in 1997 for Athena Payload;
- a *Category One* instrument by MSL payload reviews;
- Stand-alone tests in 2012 and autonomously operated by Zoë rover (50 km) in 2013 in Atacama Desert;
- CIRS (integrated optics, no optical fiber) fund by MatISSE in 2013;
- CIRS is proposed for Mars 2020, with extensive engineering developments in 2013 at JPL.

**MMRS  
(Athena,  
PIDDP,  
MIDP,  
ASTEP)**

**50 km travel in Atacama  
Desert, June 2013  
MMRS robustly performed**



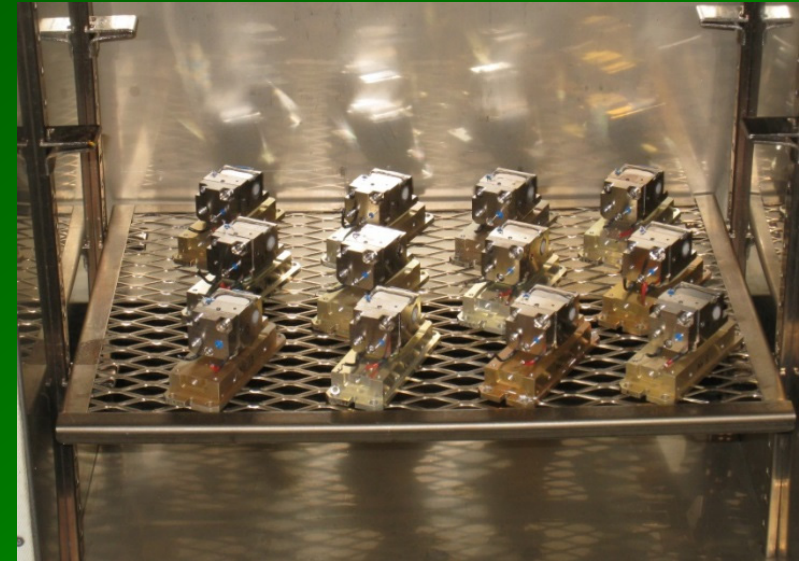
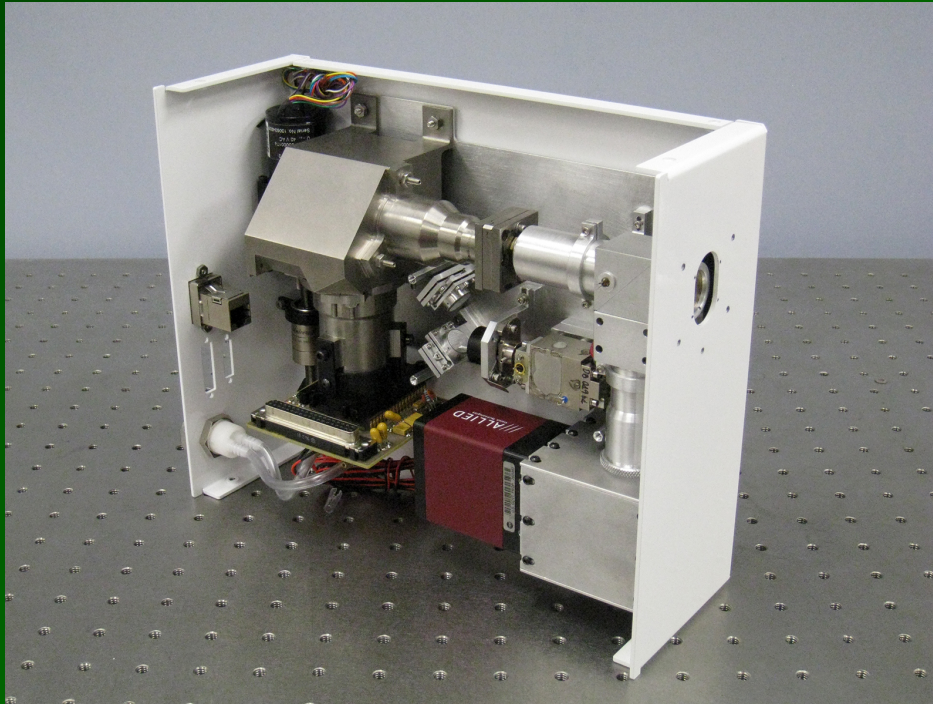
**Zoë rover**



**1 meter  
drill**

**MMRS  
probe**

# CIRS prototype (MatISSE 2013-2016)



**Laser units for testing**

**Table E.1.8-1: CIRS Key Technology Readiness Levels**

Subsystem or Component	TRL	Justification
532 nm Laser	5	MatISSE to Date
Optical Configuration	5	MatISSE to Date
E2V 47-20 Focal Plane	9	13 different missions for 47 series
KAI 2020 / M50 Focal Plane	9/8	MSL, MAHLI, Mastcam, MARDI, OSIRIS-REx
Lightfield Image Processing	6	Pipeline similarity to commercial product

# Laser Raman Spectroscopy for Planetary Explorations

## Targets -- molecular species (ID & Quantify):

- Inorganic (minerals);
- Organic (biota, biomarkers);
- Gases (atmosphere);
- Fluids & frozen (dissolved species) ;

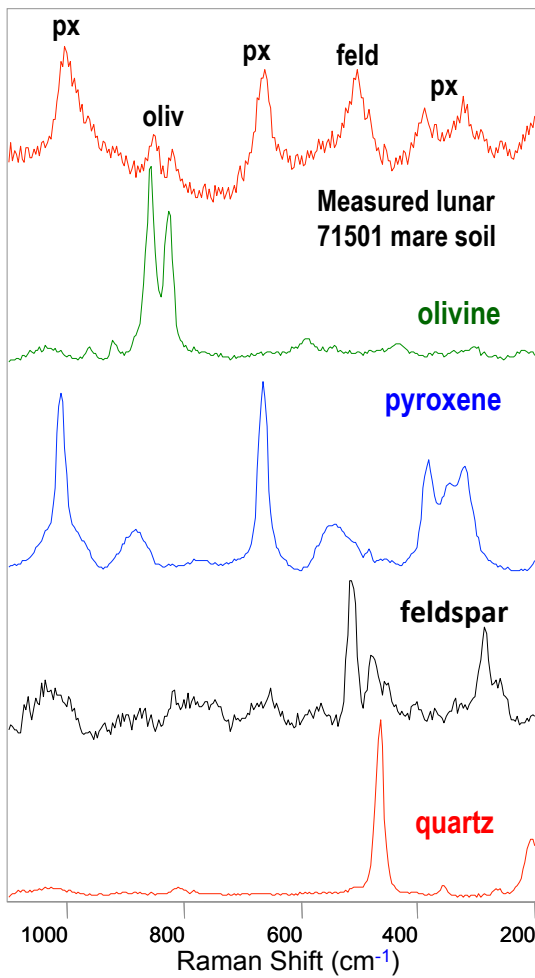
## Deployment – descending & landing missions:

- Descending through layers of atmosphere;
- on Arm & mast of rover/lander;
- Immersive probe or whole system in ocean.

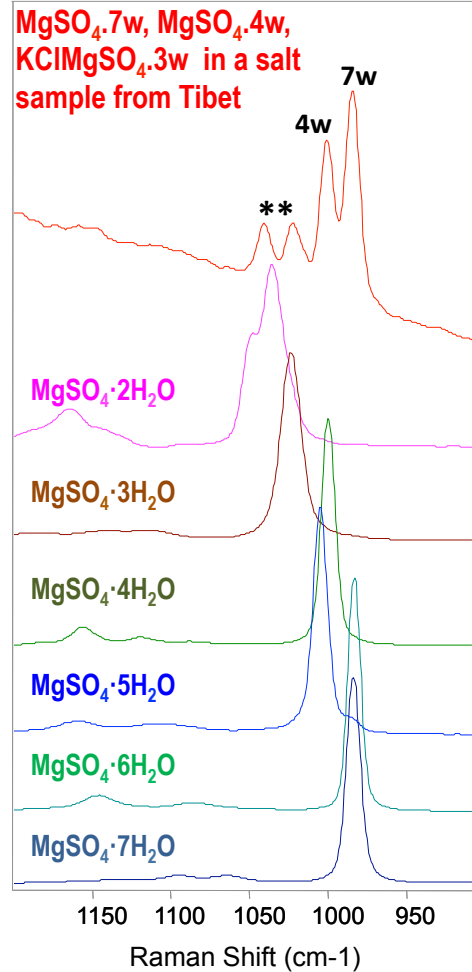
Inorganic and organic groups	covelency of major bond
carbon, sulfur, O <sub>2</sub> , N <sub>2</sub> , H <sub>2</sub>	100%
perchlorates, nitrates, carbon-(H, N)	92-95%
sulfates, carbonates	80-81%
chromates, phosphates	75-65%
borates, vanadates	59-62%
silicates	51%
(Mg, Fe, Ca)-O	19-31%

# Solid phase (ID & quantify) – Moon, Mars, asteroids

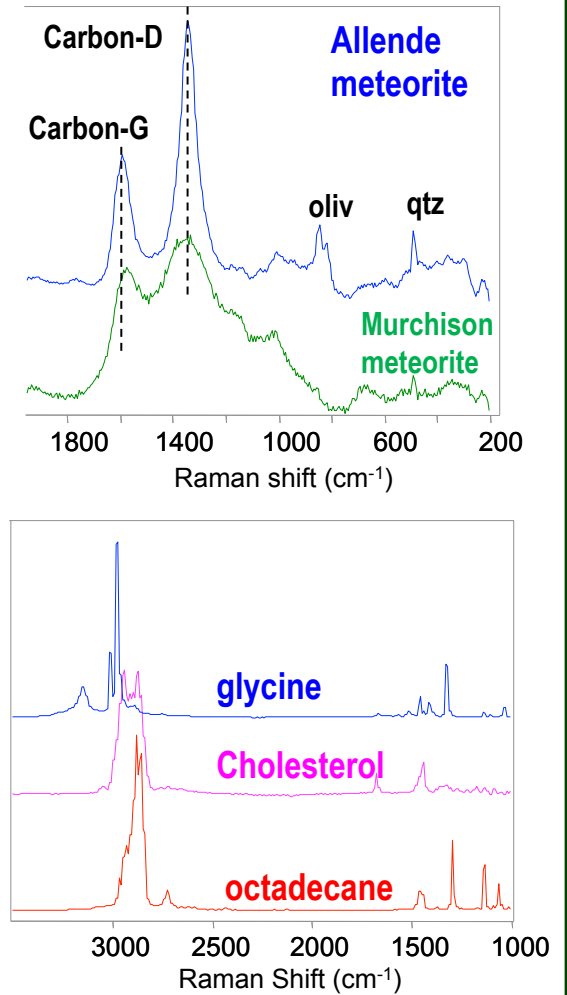
## Igneous minerals



## Hydrous salts

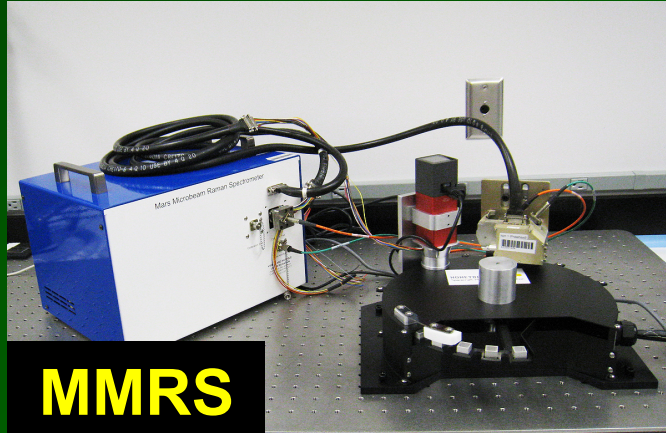


## C & organics

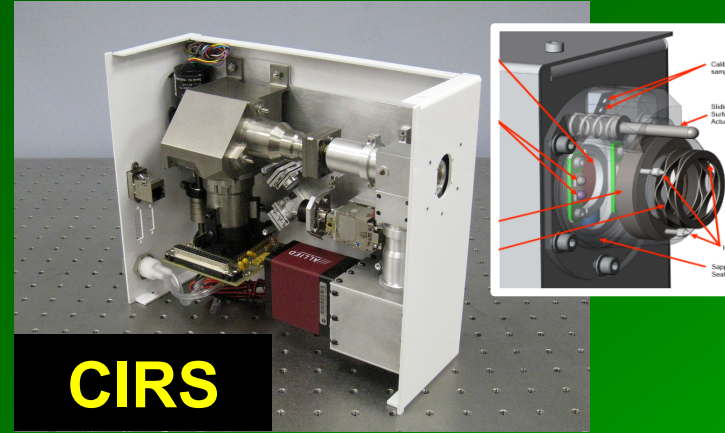


# MMRS - CIRS deployed by arm

Spectrometer-optical fiber-probe



Integrated optics + context imager



## Common features of MMRS vs. CIRS:

- Analyze the sample as is;
- ~ 27 mm working distance → detailed sensing;
- < 20  $\mu\text{m}$  laser spot → rare & trace species;
- Line-scan → Geo-context & quantification;

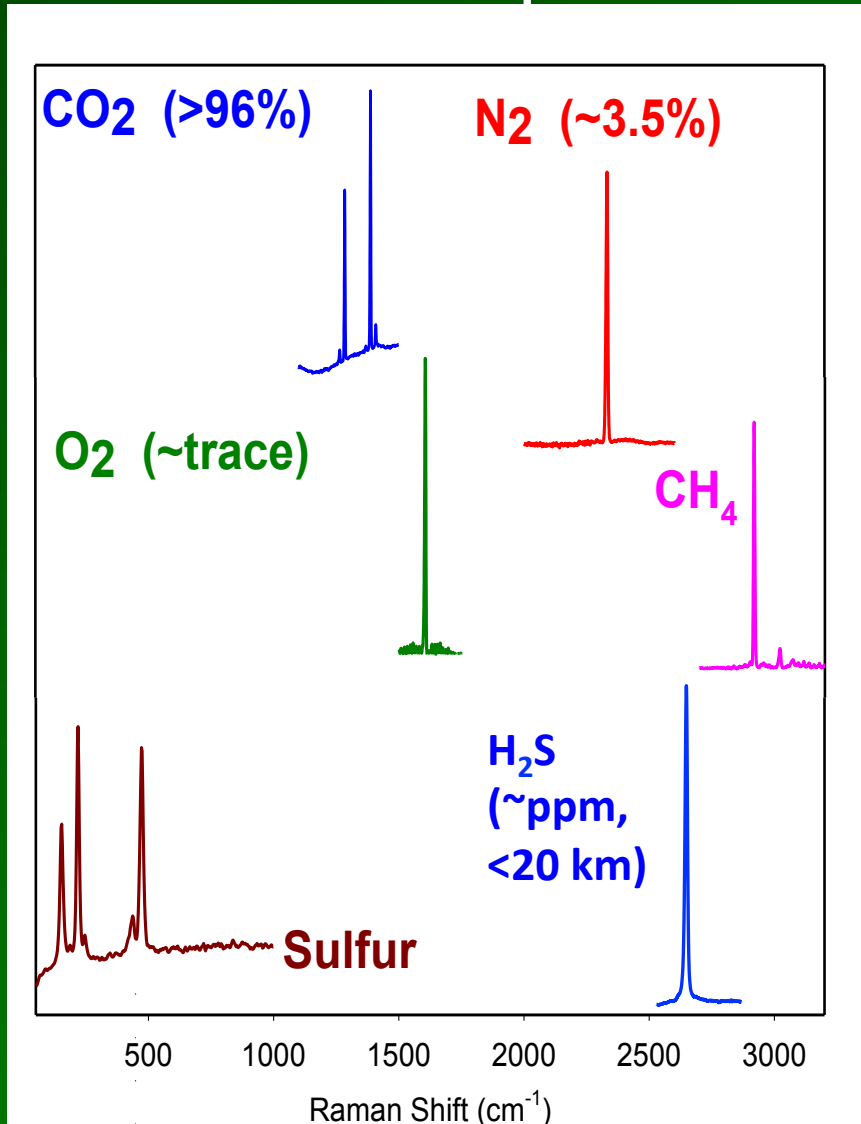
## Different features of MMRS vs. CIRS:

- MMRS has a light-mass probe, w/an optical fiber cable to be handled – for simple rover;
- CIRS has a larger-mass optics, with context imager – for heavy duty rover;

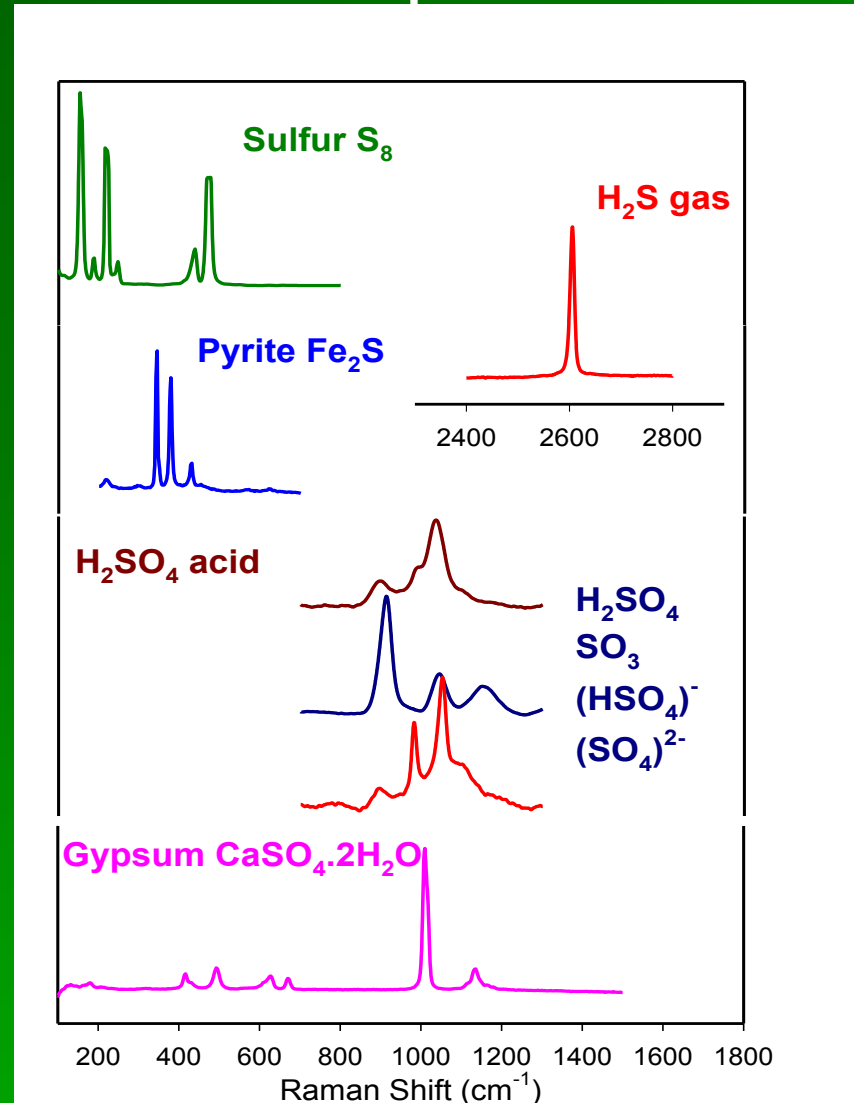
**Deployment – to be placed directly on target.**

# Gas phase (ID & quantify) – Venus etc

## Gaseous components in Venus atmosphere

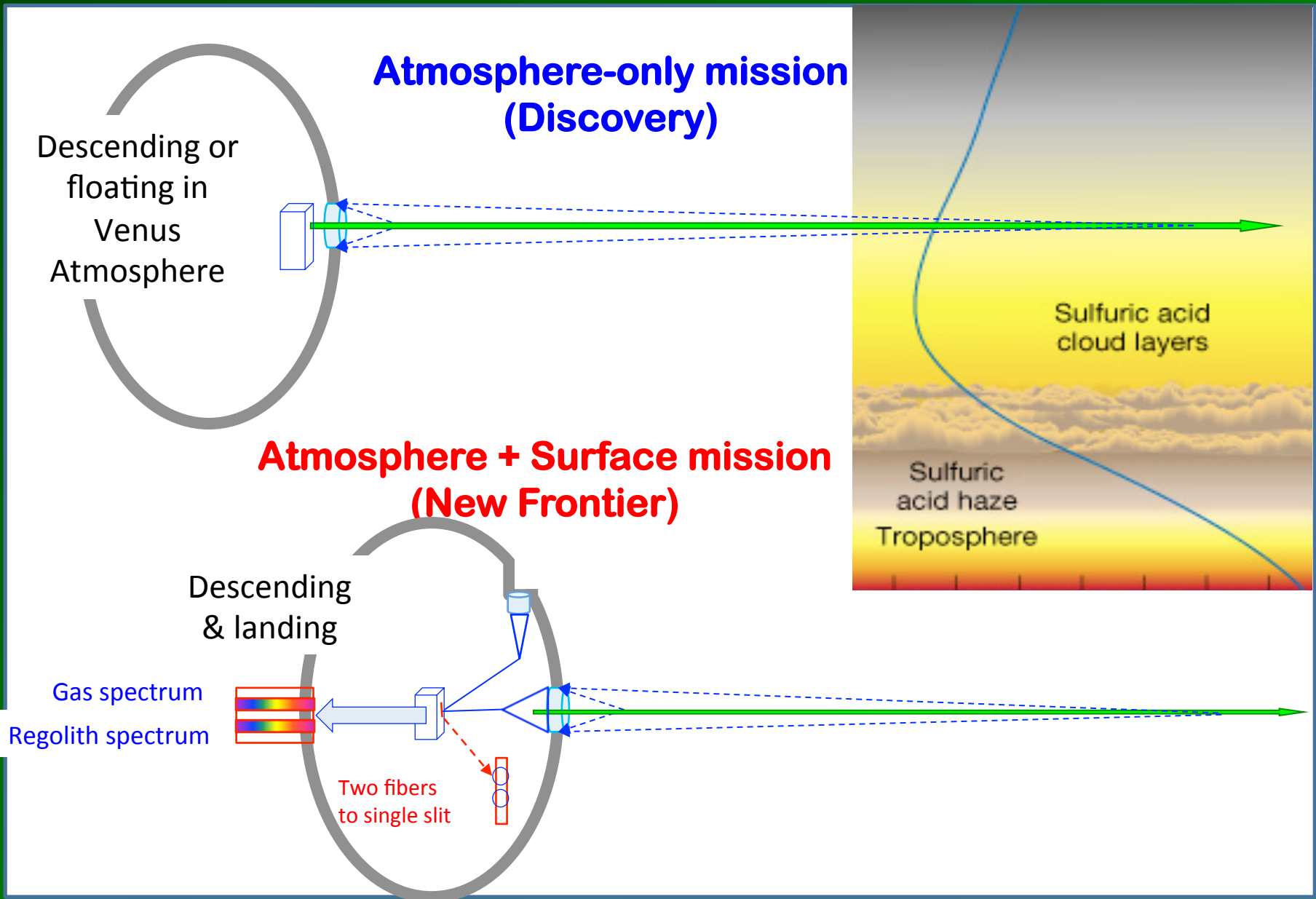


## Possible compounds in Venus atmospheric aerosol





# CIRS: flexible for Venus missions



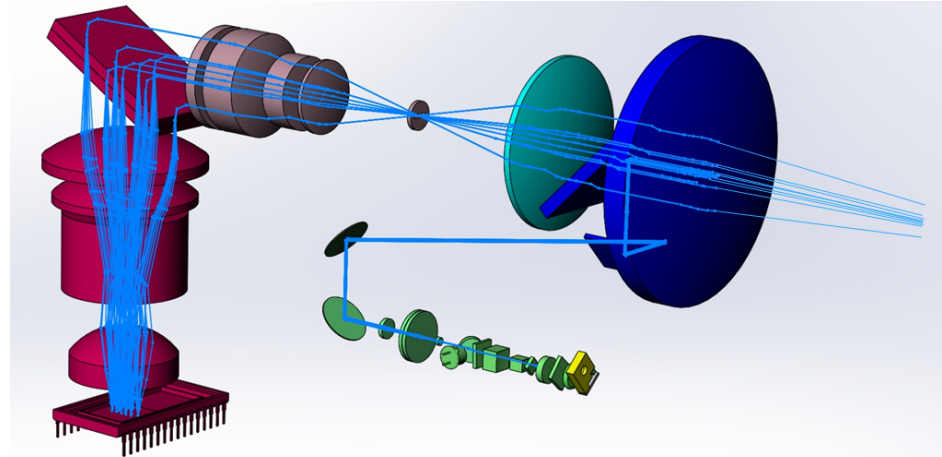
# CIRS: flexible for Venus missions



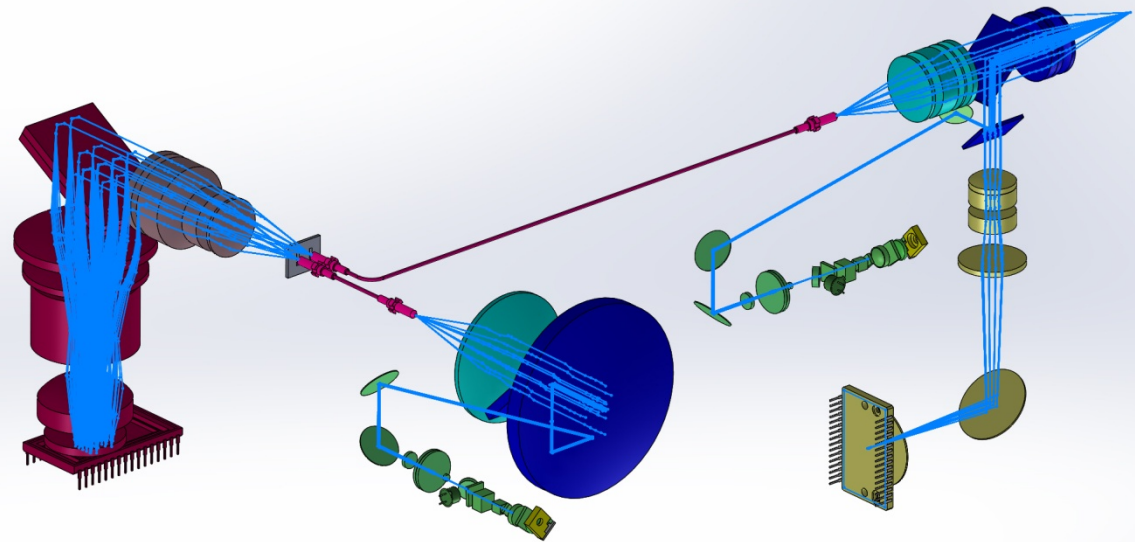
25mJ actively Q-switched  
532nm laser (MIL spec)



EMCCD provide gain and  
allow high speed  
shuttering ( $2 \mu\text{s}$  - ms)



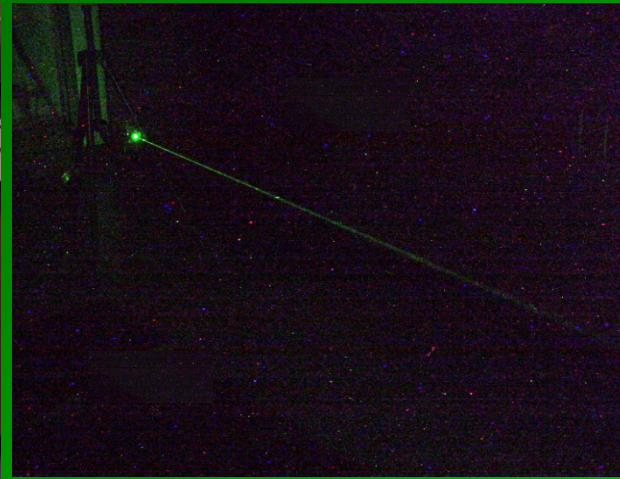
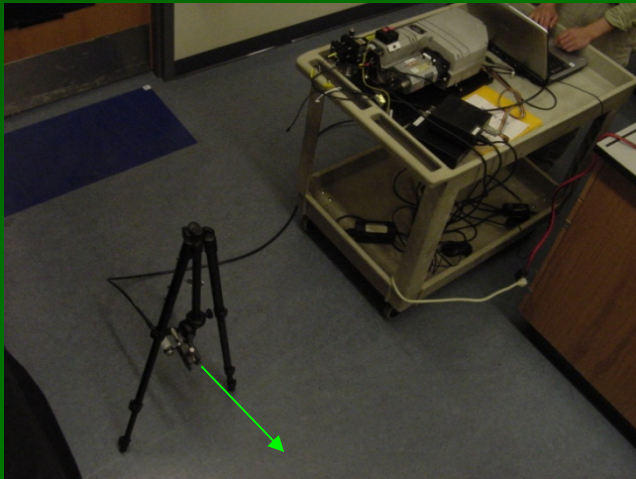
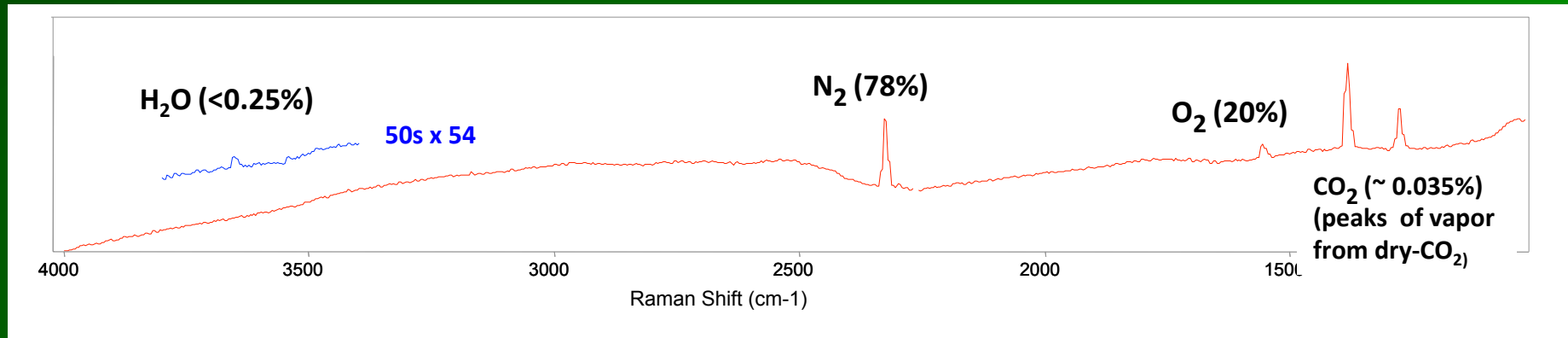
Atmosphere-only mission (Discovery)



Atmosphere & Surface mission (New Frontiers)

# A feasibility test in laboratory

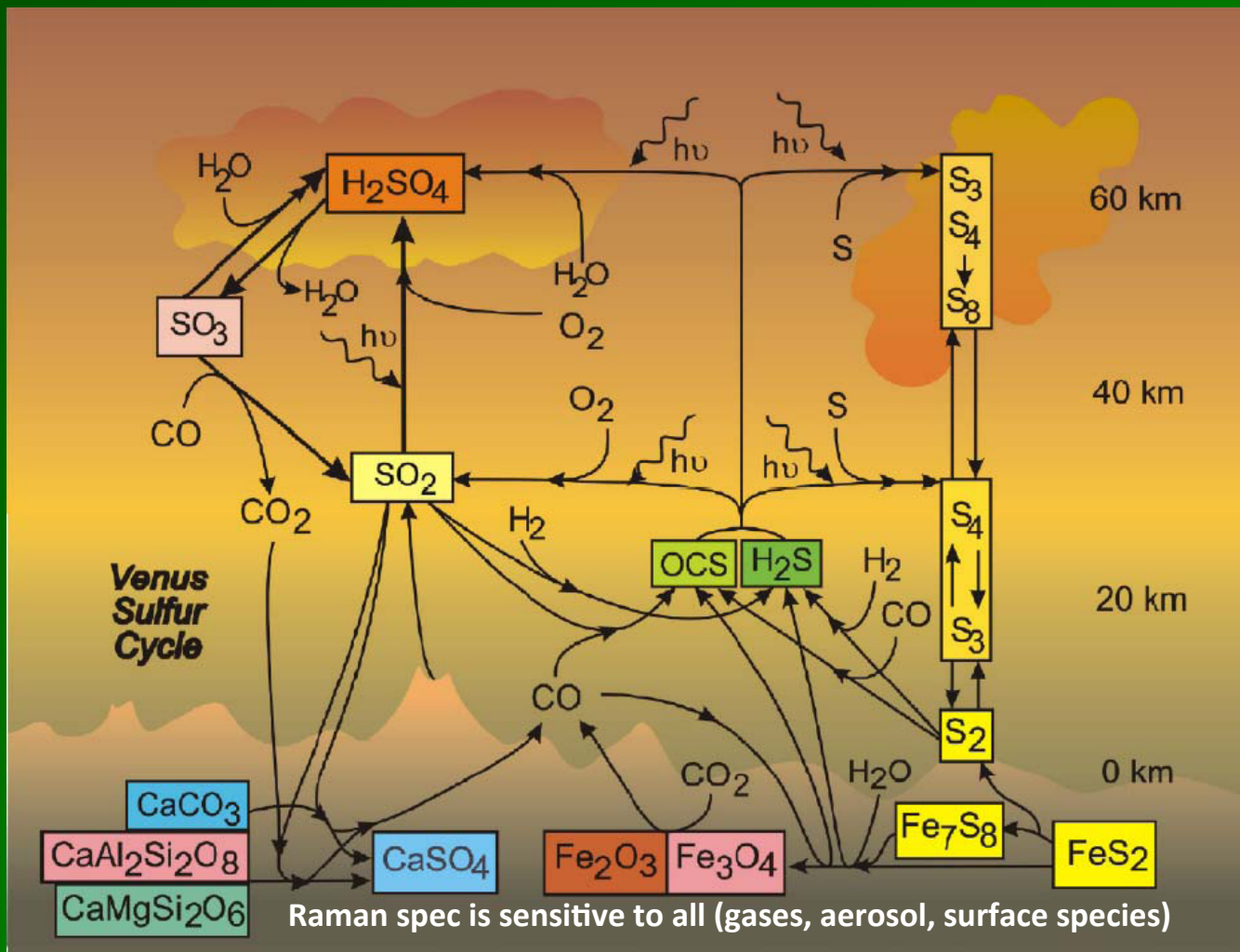
$P_{\text{Earth}} = 1 \text{ bar}$ , cw laser of 12 mw,  $\frac{1}{2}$ " optics, 5 m path:  
 $\text{N}_2$ ,  $\text{O}_2$  Raman peaks w/good SNR in 50 s



On Venus,  $\sim 92 \text{ bar}$ , cw laser of 1w, 2" optics, 5-10 m path,  
Need  $3 \times 10^{-4} \text{ s}$  to get gas peaks with similar SNR.

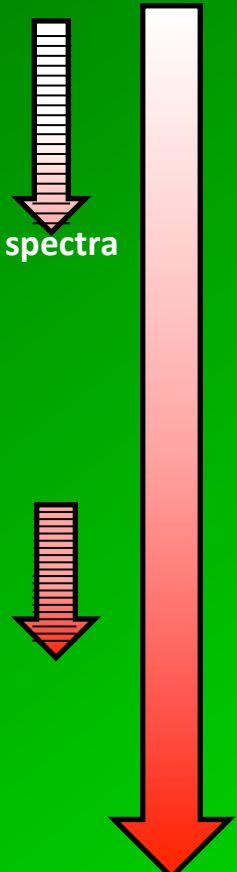
Time is a major limitation for Venus missions – we would use a meas-logic different from SAM, i.e.

**Get a Raman spectrum in every second**  
(then detailed data digestion back Earth)



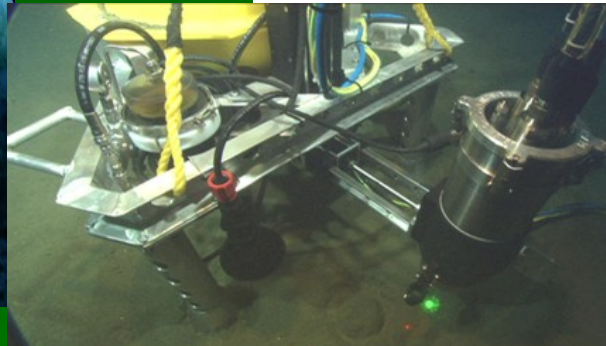
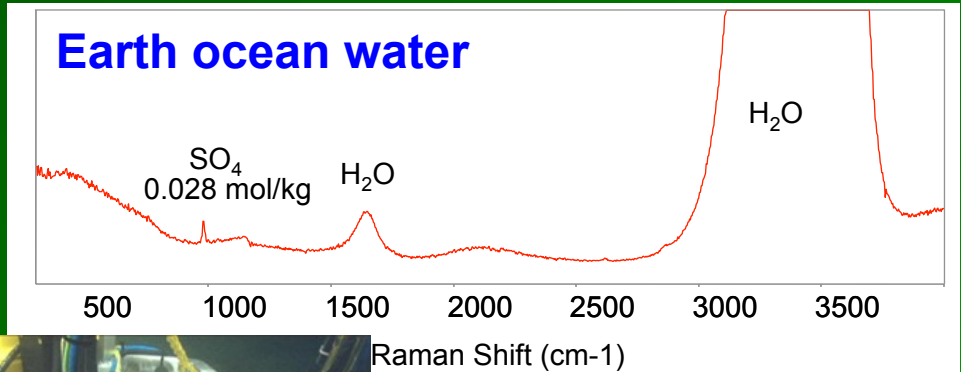
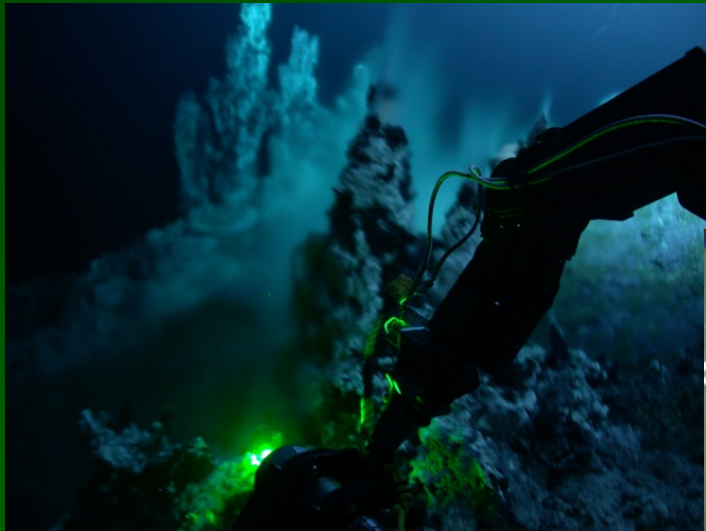
**Data density**

20 min  
= 1200 spectra

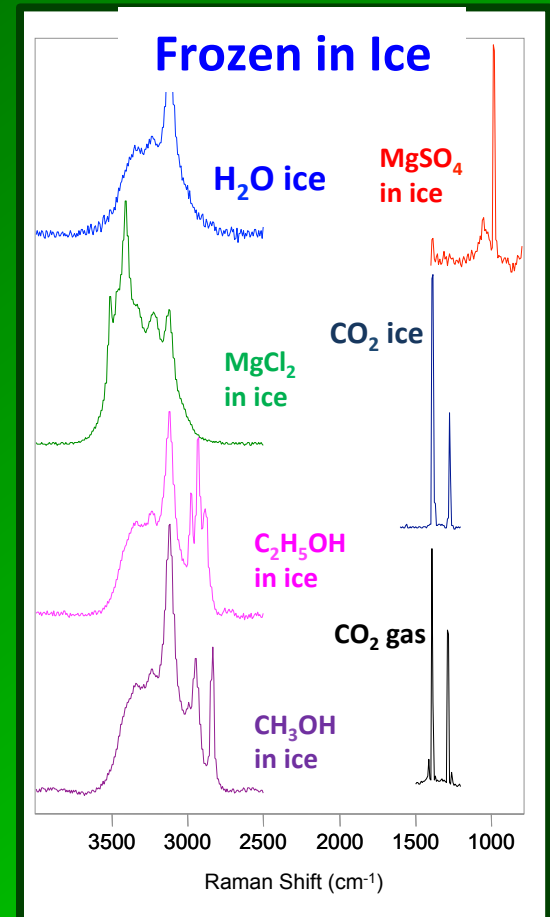
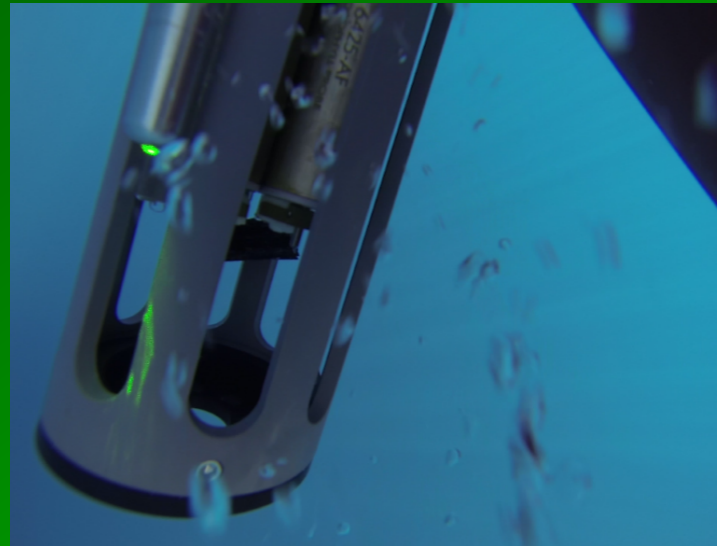


# Liquid · frozen · dissolved species

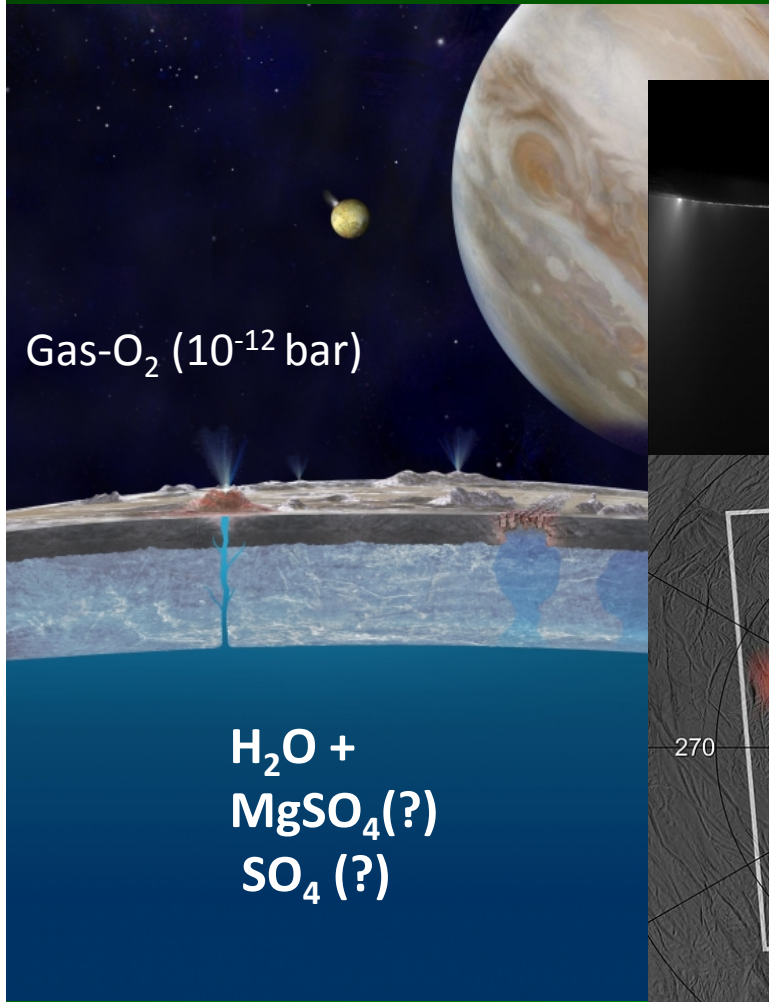
Raman system used in deep ocean



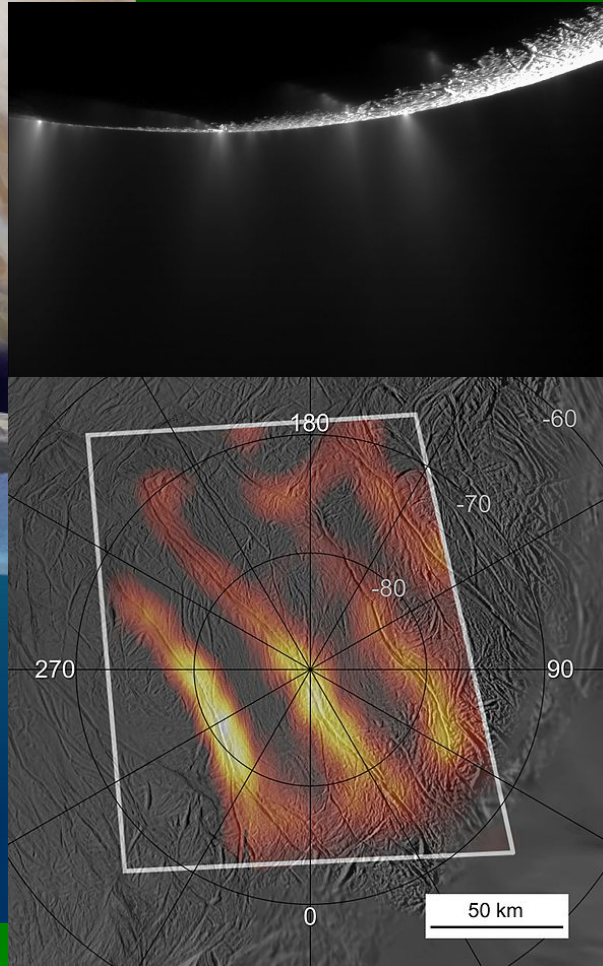
Raman probing lake water



# Exploration of icy world

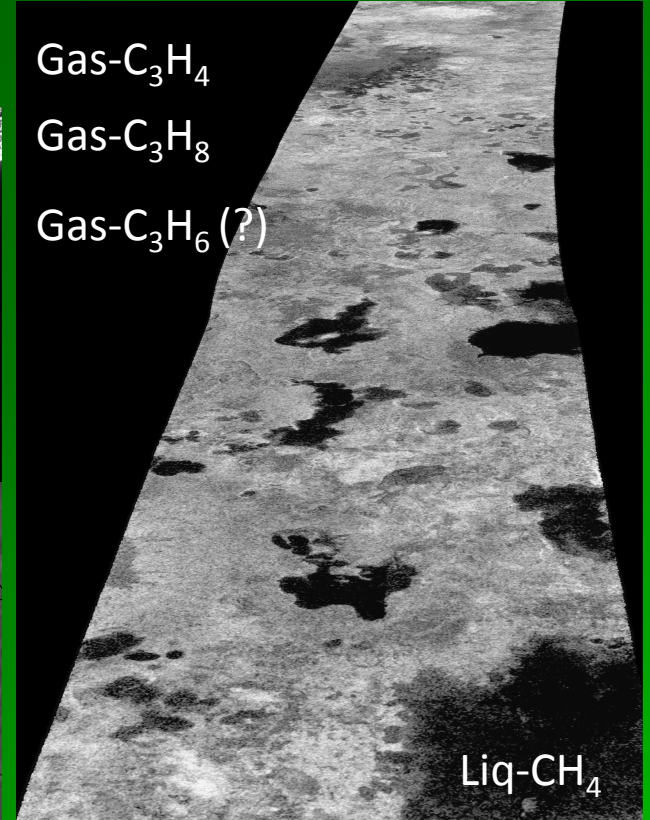


Subsurface water  
ocean of **Europa**



Subsurface salty water  
ocean of **Enceladus**  
(organics in jet dust)

Gas-C<sub>3</sub>H<sub>4</sub>  
Gas-C<sub>3</sub>H<sub>8</sub>  
Gas-C<sub>3</sub>H<sub>6</sub> (?)



hydrocarbon lakes on **Titan**  
northern hemisphere  
(2013 Cassini-Huygens)

# Acknowledgements – NASA PMD funds

